

# Heavy Quarkonium magnetic (and electric) dipole transitions in pNRQCD

Mainly based on Phys. Rev. D87, 074024. A. Pineda and J. Segovia

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# Motivation

- ▶ 1st principle computation of heavy quarkonium properties from QCD. To give model independent predictions with model independent errors.
- ▶ Determination of Standard Model parameters:  $m_Q$ ,  $\alpha_s$ , ...

Potential Non-Relativistic QCD in the weak coupling regime is ideal for this.  
 $m \gg mv \gg mv^2$

$$\left. \begin{array}{l} \left( i\partial_0 - \frac{\mathbf{p}^2}{2m} - V_s^{(0)}(r) \right) \Phi(\mathbf{r}) = 0 \\ + \text{corrections to the potential} \\ + \text{interaction with other low} \\ \quad \text{energy degrees of freedom} \end{array} \right\} \text{pNRQCD (Pineda, Soto)} \quad E \sim mv^2$$

In the strict weak coupling regime the starting point is

$$V_s^{(0)} \simeq V^C \equiv -C_f \frac{\alpha_s(\mu)}{r}.$$

To which extent is suitable a weak coupling description of heavy quarkonium?

$$V_s^{(0)} = -C_f \frac{\alpha_V(1/r)}{r} = -C_f \frac{\alpha_s(\mu)}{r} \left( 1 + a_1 \frac{\alpha_s(\mu)}{4\pi} + a_2 \frac{\alpha_s^2(\mu)}{16\pi^2} + \dots \right)$$

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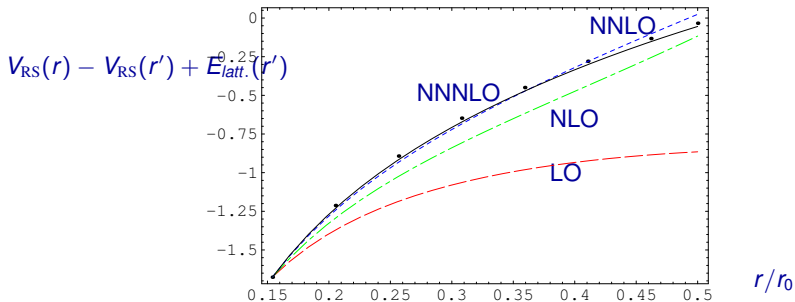
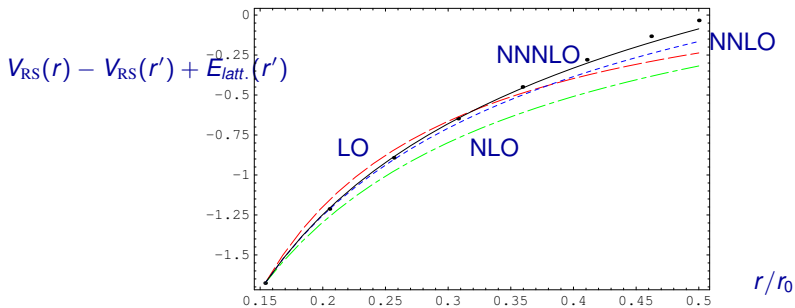
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## Proposal: Reorganization of perturbation theory

Strict perturbation theory

$$H^{(0)} = \frac{\mathbf{p}^2}{m} + V^C \longrightarrow E_n^C, \phi_n^C(\mathbf{r})$$

Coulomb potential

$$V_s^C = -G_f \frac{\alpha_s(\mu)}{r}$$

Relativistic +  $\delta V^{(0)}$  corrections:

$$\Delta H = \delta V^{(0)} + \frac{V^{(1)}}{m} + \frac{V^{(2)}}{m^2} + \dots$$



## Proposal: Reorganization of perturbation theory

Improved perturbation theory "acceleration of perturbation theory"

$$H^{(0)} = \frac{\mathbf{p}^2}{m} + V^{(0)} \longrightarrow E_n^{(0)}, \phi_n^{(0)}(\mathbf{r})$$

Keep the static potential exactly

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Relativistic corrections:

$$\Delta H = \frac{V^{(1)}}{m} + \frac{V^{(2)}}{m^2} + \dots$$

Applied to inclusive electromagnetic decay ratios (Kiyo, Pineda, Signer).  
Apply it to M1 radiative transitions.

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## Allowed transitions

Strict weak coupling: Brambilla, Jia, Vairo

$$\Gamma(n^3S_1 \rightarrow n^1S_0\gamma) = \frac{4}{3}\alpha e_Q^2 \frac{k_\gamma^3}{m^2} \left[ 1 + 2\kappa - \frac{5}{3} \frac{\langle p^2 \rangle_n^C}{m^2} \right],$$

$$\Gamma(n^3P_J \rightarrow n^1P_1\gamma) = \frac{3\Gamma(n^1P_1 \rightarrow n^3P_J\gamma)}{2J+1} = \frac{4}{3}\alpha e_Q^2 \frac{k_\gamma^3}{m^2} \left[ 1 + 2\kappa - d_J \frac{\langle p^2 \rangle_{n1}^C}{m^2} \right],$$

where  $d_0 = 1$ ,  $d_1 = 2$ ,  $d_2 = 8/5$ ,

$$k_\gamma = |\vec{k}| = \frac{M_H^2 - M_{H'}^2}{2M_H},$$

and the anomalous magnetic moment of the heavy quark reads

$$\kappa = \kappa^{(1)}\alpha_s(m)$$

Leading order= potential models

## Allowed transitions

Improved weak coupling: Pineda, Segovia

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and the anomalous magnetic moment of the heavy quark reads

$$\kappa = \kappa^{(1)}\alpha_s(m) + \kappa^{(2)}\alpha_s^2(m)$$

We also implement the renormalon.

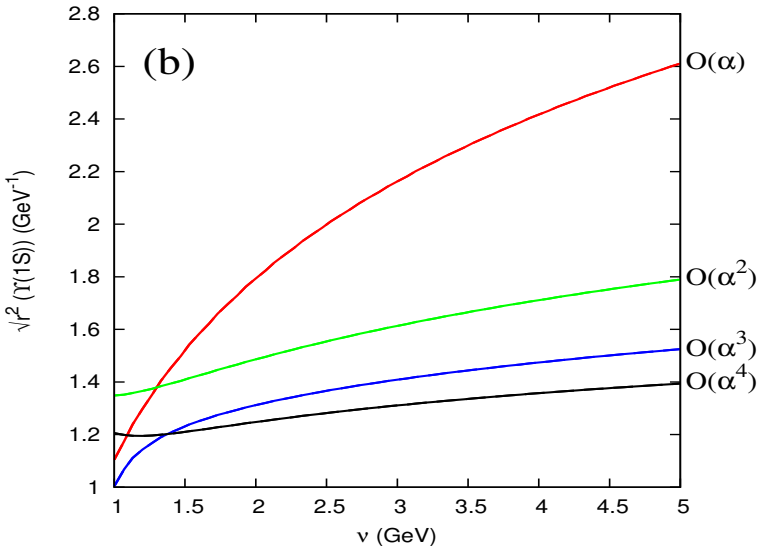


Figure :  $\sqrt{\langle r^2 \rangle_{10}}$  using the static potential  $V_{RS'}^{(N)}$  at different orders in perturbation theory:  $N = 0, 1, 2, 3$ . Dashed lines with  $\nu_f = 0$ . Continuous lines with  $\nu_f = 0.7$  GeV. In both cases  $\nu_r = \infty$  GeV.

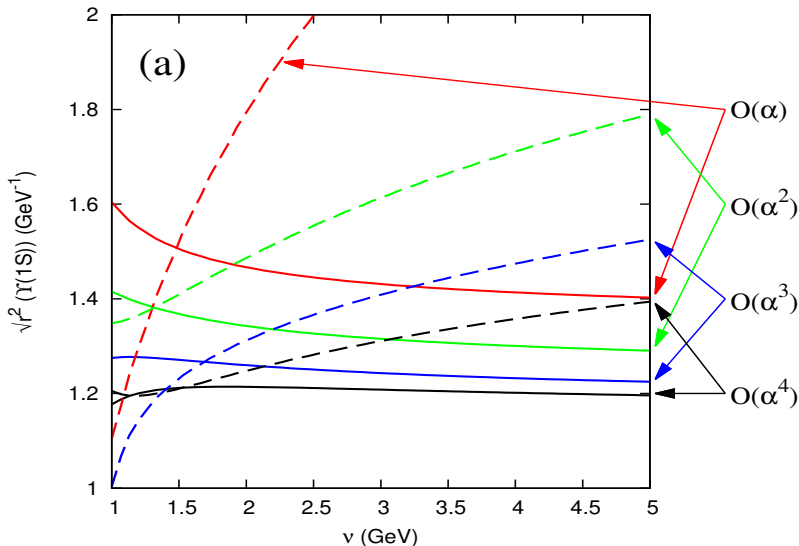


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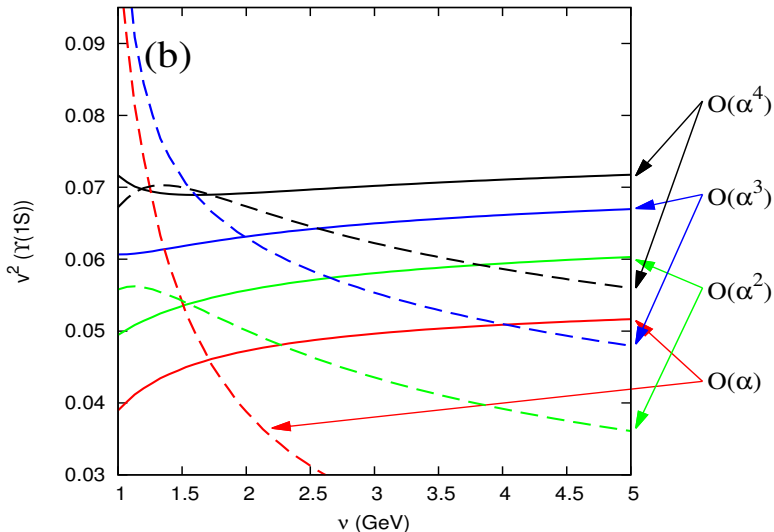


Figure :  $v_{10}^2$  using the static potential  $V_{RS'}^{(N)}$  at different orders in perturbation theory:  $N = 0, 1, 2, 3$ . Dashed lines with  $\nu_r = \infty$ . Continuous lines with  $\nu_r = 0.7$  GeV. In both cases  $\nu_f = 0.7$  GeV.



|                     | LO    | $\mathcal{O}(\alpha_s)$ | $\mathcal{O}(\alpha_s^2)$ | $\mathcal{O}(v^2)$ | $\alpha_s \times \mathcal{O}(\alpha_s^2)$ | $v \times \mathcal{O}(v^2)$ |
|---------------------|-------|-------------------------|---------------------------|--------------------|---|-----------------------------|
| $\delta\Gamma$ (eV) | 14.87 | 1.29                    | 0.73                      | -1.71              | 0.15                                      | -0.45                       |

Table : *The leading and subleading contributions to  $\Gamma_{\Upsilon(1S) \rightarrow \eta_b(1S)\gamma}$ . The last two numbers are error estimates obtained by multiplying the subleading  $\mathcal{O}(\alpha_s^2)$  contribution by  $\alpha_s$  and the subleading  $\mathcal{O}(v^2)$  contribution by  $v$ .*

$$\Gamma_{\Upsilon(1S) \rightarrow \eta_b(1S)\gamma} = 15.18 \pm 0.45 (\mathcal{O}(v^3))_{-0.05}^{-0.12} (N_m)_{+0.03}^{-0.04} (\alpha_s)_{+0.20}^{-0.20} (m_{\overline{\text{MS}}}) \text{ eV},$$

which after combining the errors in quadrature reads

$$\Gamma_{\Upsilon(1S) \rightarrow \eta_b(1S)\gamma} = 15.18(51) \text{ eV}.$$

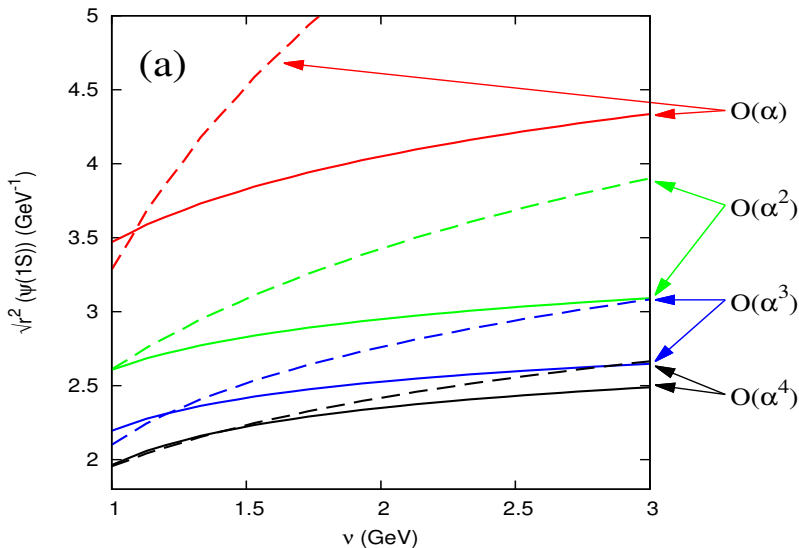


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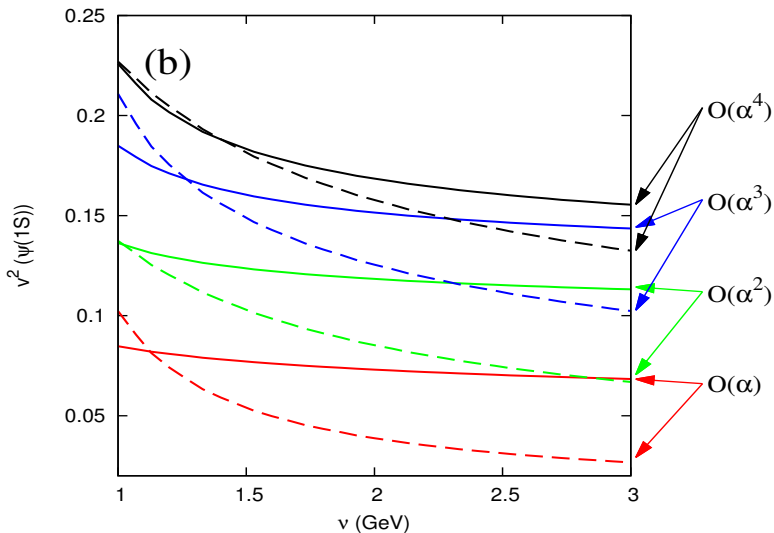


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|                      | LO   | $\mathcal{O}(\alpha_s)$ | $\mathcal{O}(\alpha_s^2)$ | $\mathcal{O}(v^2)$ | $\alpha_s \times \mathcal{O}(\alpha_s^2)$ | $v \times \mathcal{O}(v^2)$ |
|----------------------|------|-------------------------|---------------------------|--------------------|---|-----------------------------|
| $\delta\Gamma$ (keV) | 2.34 | 0.33                    | 0.16                      | -0.71              | 0.05                                      | -0.30                       |

Table : *The leading and subleading contributions to  $\Gamma_{J/\psi(1S)\rightarrow\eta_c(1S)\gamma}$ . The last two numbers are error estimates obtained by multiplying the subleading  $\mathcal{O}(\alpha_s^2)$  contribution by  $\alpha_s$  and the subleading  $\mathcal{O}(v^2)$  contribution by  $v$ .*

$$\Gamma_{J/\psi(1S)\rightarrow\eta_c(1S)\gamma} = 2.12 \pm 0.30(\mathcal{O}(v^3))_{-0.23}^{+0.21}(N_m)_{+0.02}^{-0.02}(\alpha_s)_{+0.11}^{-0.10}(m_{\overline{MS}}) \text{ keV},$$

which, after combining the errors in quadrature, reads

$$\Gamma_{J/\psi(1S)\rightarrow\eta_c(1S)\gamma} = 2.12(40) \text{ keV}.$$

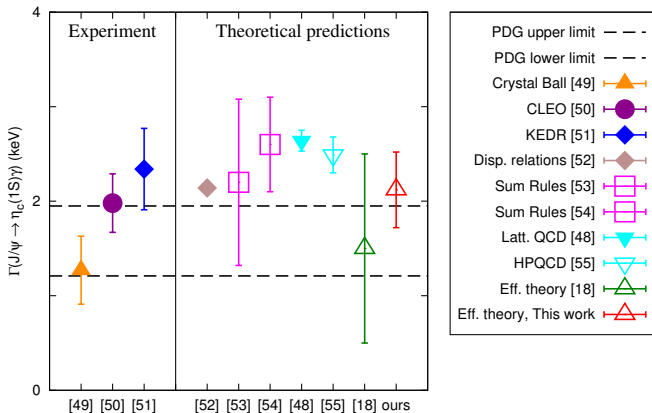


Figure : Comparison of different theoretical and experimental predictions for  $\Gamma_{J/\psi \rightarrow \eta_c \gamma}$ .

|   | $b\bar{b}(1S)$ | $c\bar{c}(1S)$ | $b\bar{b}(1P)$ | $b\bar{b}(2S)$ |
|---|----------------|----------------|----------------|----------------|
| $v$   | 0.26           | 0.43           | 0.25           | 0.24           |
| $\sqrt{\langle r^2 \rangle}(\text{GeV}^{-1})$ | 1.2            | 2.2            | 2.1            | 2.9            |

Table : Estimates for  $v \equiv \sqrt{\langle p^2 \rangle}/m^2$  and  $\sqrt{\langle r^2 \rangle}$  for the heavy quarkonium states. For the  $b\bar{b}(2S)$  state the number we give for  $v$  is quite uncertain.

## Hindered transitions. Strict weak coupling: Brambilla, Jia, Vairo

$$\Gamma(n^3 S_1 \rightarrow n^1 S_0 \gamma) \stackrel{n \neq n'}{\equiv} \frac{4}{3} \alpha e_Q^2 \frac{k_\gamma^3}{m^2} \left[ \frac{k_\gamma^2}{24} n' \langle r^2 \rangle_n^C + \frac{5}{6} \frac{n' \langle p^2 \rangle_n^C}{m^2} - \frac{2}{m^2} \frac{n' \langle V_{S^2}^C(\vec{r}) \rangle_n}{E_n^C - E_{n'}^C} \right]^2,$$

$$\Gamma(n^1 S_0 \rightarrow n^3 S_1 \gamma) \stackrel{n \neq n'}{\equiv} 4 \alpha e_Q^2 \frac{k_\gamma^3}{m^2} \left[ \frac{k_\gamma^2}{24} n' \langle r^2 \rangle_n^C + \frac{5}{6} \frac{n' \langle p^2 \rangle_n^C}{m^2} + \frac{2}{m^2} \frac{n' \langle V_{S^2}^C(\vec{r}) \rangle_n^C}{E_n^C - E_{n'}^C} \right]^2,$$

$$V_{S^2}(\vec{r}) = \frac{4}{3} \pi C_f D_{S^2, s}^{(2)}(\nu) \delta^{(3)}(\vec{r})$$

$$D_{S^2, s}^{(2)}(\nu) = \alpha_s(\nu)$$

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$$\Gamma(n^1 S_0 \rightarrow n'^3 S_1 \gamma) \stackrel{n \neq n'}{=} 4 \alpha e_Q^2 \frac{k_\gamma^3}{m^2} \left[ \frac{k_\gamma^2}{24} n' \langle r^2 \rangle_n + \frac{5}{6} \frac{n' \langle p^2 \rangle_n}{m^2} + \frac{2}{m^2} \frac{n' \langle V_{S^2}(\vec{r}) \rangle_n}{E_n - E_{n'}} \right]^2,$$

$$V_{S^2}(\vec{r}) = \frac{4}{3} \pi C_f D_{S^2, s}^{(2)}(\nu) \delta^{(3)}(\vec{r})$$

$$D_{S^2, s}^{(2)}(\nu) = \alpha_s(\nu) c_F^2(\nu) - \frac{3}{2\pi C_f} (d_{sv}(\nu) + C_f d_{vv}(\nu))$$

depends on the NRQCD Wilson coefficients. With LL accuracy they read

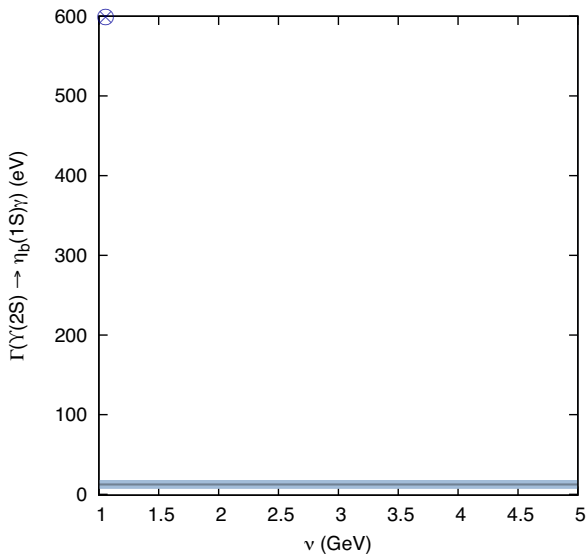
$$c_F(\nu) = z^{-C_A}, \quad d_{sv}(\nu) = d_{sv}(m),$$

$$d_{vv}(\nu) = d_{vv}(m) + \frac{C_A}{\beta_0 - 2C_A} \pi \alpha_s(m) (z^{\beta_0 - 2C_A} - 1),$$

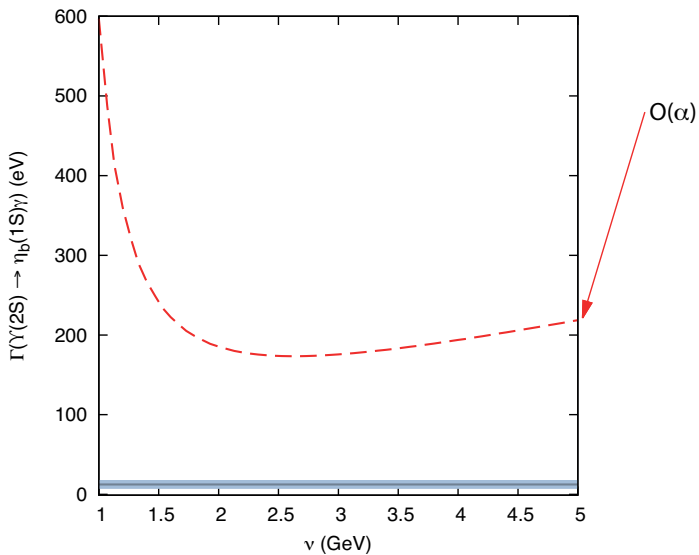
$$z = \left[ \frac{\alpha_s(\nu)}{\alpha_s(m)} \right]^{\frac{1}{\beta_0}} \simeq 1 - \frac{1}{2\pi} \alpha_s(\nu) \ln \left( \frac{\nu}{m} \right),$$

$$d_{sv}(m) = C_f \left( C_f - \frac{C_A}{2} \right) \pi \alpha_s(m), \quad d_{vv}(m) = - \left( C_f - \frac{C_A}{2} \right) \pi \alpha_s(m).$$

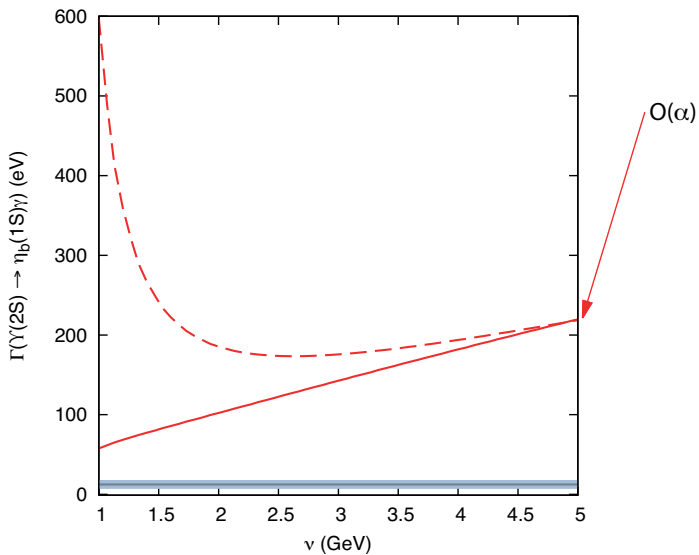




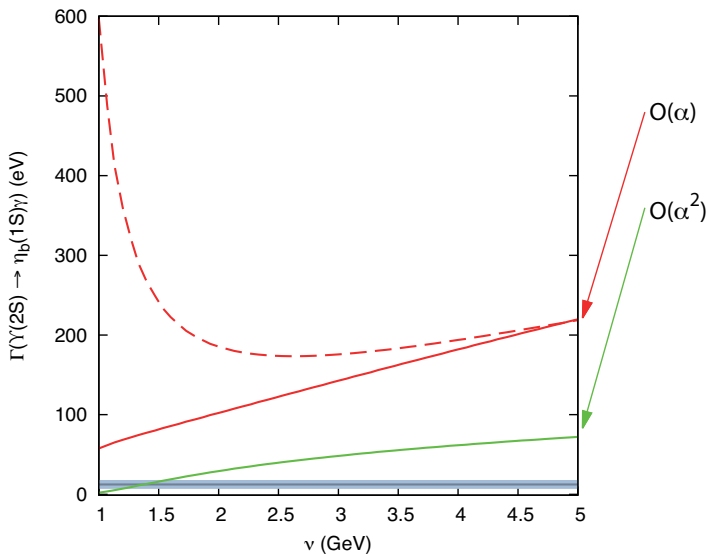
no RG ,  $V_s^{(0)} = -C_f \frac{\alpha_s(\mu)}{r}$  ,  $\mu = 1 \text{ GeV}$



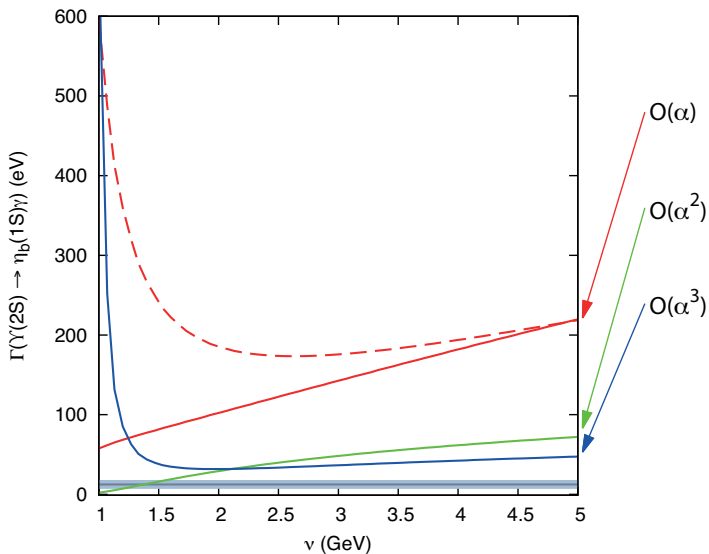
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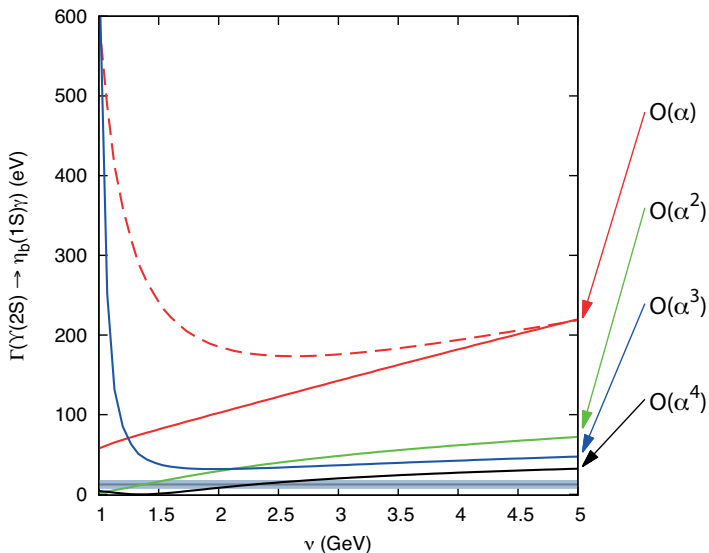
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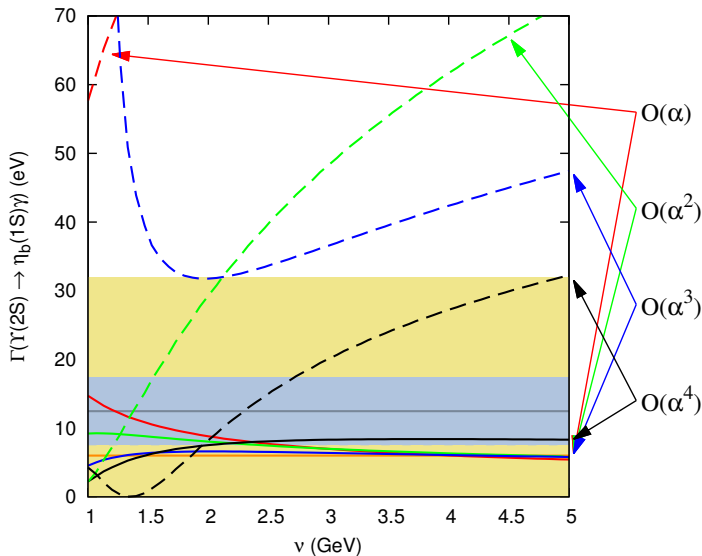
$$RG, V_s^{(0)} = -C_f \frac{\alpha_s(\mu)}{r} \left( 1 + a_1 \frac{\alpha_s(\mu)}{4\pi} \right)$$



$$RG, V_s^{(0)} = -C_f \frac{\alpha_s(\mu)}{r} \left( 1 + a_1 \frac{\alpha_s(\mu)}{4\pi} + a_2 \frac{\alpha_s^2(\mu)}{(4\pi)^2} \right)$$



$$RG, V_s^{(0)} = -C_f \frac{\alpha_s(\mu)}{r} \left( 1 + a_1 \frac{\alpha_s(\mu)}{4\pi} + a_2 \frac{\alpha_s^2(\mu)}{(4\pi)^2} + a_3 \frac{\alpha_s^3(\mu)}{(4\pi)^3} \right)$$



$$\text{RG} , V_s^{(0)} = -C_f \frac{\alpha_s(1/r)}{r} \left( 1 + a_1 \frac{\alpha_s(1/r)}{4\pi} + a_2 \frac{\alpha_s^2(1/r)}{(4\pi)^2} + a_3 \frac{\alpha_s^3(1/r)}{(4\pi)^3} \right)$$

| Prefactor (keV) | $\mathcal{A}(r^2)$ | $\mathcal{A}(\vec{p}^2)$ | $\mathcal{A}(V_{S^2})$ | $\Gamma$ (eV) |
|-----------------|--------------------|--------------------------|------------------------|---------------|
| 10.3342         | 0.022              | 0.039                    | -0.042                 | 6.3           |

Table : *The prefactor, the terms inside the brackets, and the total decay width*

$\Gamma_{\Upsilon(2S) \rightarrow \eta_b(1S)\gamma}$ .

$$\Gamma_{\Upsilon(2S) \rightarrow \eta_b(1S)\gamma}^{(\text{th})} = 0.006 \pm 0.006 (\mathcal{O}(v^5))_{-0.006}^{+0.026} (N_m)_{+0.001}^{-0.001} (\alpha_s)_{+0.000}^{-0.000} (m_{\overline{MS}}) \text{ keV}.$$

$$\Gamma_{\Upsilon(2S) \rightarrow \eta_b(1S)\gamma}^{(\text{th})} = 6_{-06}^{+26} \text{ eV}.$$

Matrix element.

Experimental number: 0.035(7)

Ours:  $0.025_{-0.025}^{+0.031}$

Lewis, Woloshyn:  $\mathcal{O}(v^4) = 0.080(5)$ ;  $\mathcal{O}(v^6) = 0.032(5)$

Both agreement but different physics. To be clarified.



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Experimental number: 0.035(7)

Ours:  $0.025_{-0.025}^{+0.031}$

Lewis, Woloshyn:  $\mathcal{O}(v^4) = 0.080(5)$ ;  $\mathcal{O}(v^6) = 0.032(5)$

Both agreement but different physics. To be clarified.

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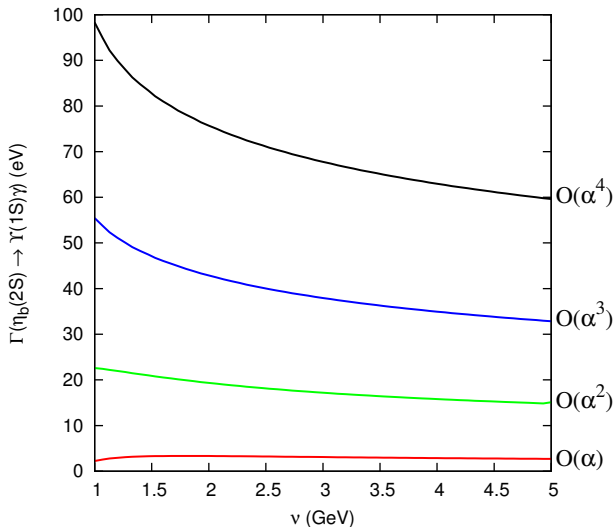


Figure : Plot of  $\Gamma_{\eta_b(2S) \rightarrow \Upsilon(1S)\gamma}$  using the static potential  $V_{RS'}^{(N)}$  at different orders in perturbation theory:  $N = 0, 1, 2, 3$  with  $\nu_r = \nu_f = 0.7$  GeV.

## E1 transitions

Brambilla, Pietrulewicz and Vairo

$$\Gamma_{n^3P_J \rightarrow n'^3S_1\gamma} = \frac{4}{9} \alpha_{em} e_Q^2 k_\gamma^3 I_3^2(n1 \rightarrow n'0) \\ \times \left( 1 + R - \frac{k_\gamma^2}{60} \frac{I_5}{I_3} - \frac{k_\gamma}{6m} + \frac{\kappa k_\gamma}{2m} \left[ \frac{J(J+1)}{2} - 2 \right] \right),$$

where

$$I_N \equiv \int_0^\infty dr r^N R_{n'0}(r) R_{n1}(r).$$

$R$  contains all of the wave-function corrections due to higher-order potentials, the relativistic correction of the kinetic energy,  $-\mathbf{p}^4/4m^3$ , and higher-order Fock state contributions due to intermediate color-octet states.



| process  | $\Gamma_{\text{pNRQCD}}^{\text{LO}}/\text{keV}$ | $\Gamma_{\text{pNRQCD}}^{\text{NLO}}/\text{keV}$ | $\Gamma_{\text{mod}}/\text{keV}$ | $\Gamma_{\text{exp}}^{\text{PDG}}/\text{keV}$ |
|--|---|--|----------------------------------|---|
| $\chi_{b0}(1P) \rightarrow \Upsilon(1S)\gamma$ | 31.8  | $29.7 \pm 3.1$                                   | 25.7-27.0                        | -   |
| $\chi_{b1}(1P) \rightarrow \Upsilon(1S)\gamma$ | 40.3  | $35.8 \pm 4.0$                                   | 29.8-31.2                        | -   |
| $\chi_{b2}(1P) \rightarrow \Upsilon(1S)\gamma$ | 45.9  | $40.6 \pm 4.6$                                   | 33.0-34.2                        | -   |
| $h_b(1P) \rightarrow \eta_b(1S)\gamma$         | 60.8  | $44.3 \pm 6.1$                                   | -                                | -   |
| $\Upsilon(2S) \rightarrow \chi_{b0}(1P)\gamma$ | 1.52  | $1.13 \pm 0.15$                                  | 0.72-0.73                        | $1.22 \pm 0.16$                               |
| $\Upsilon(2S) \rightarrow \chi_{b1}(1P)\gamma$ | 2.26  | $1.94 \pm 0.23$                                  | 1.62-1.65                        | $2.21 \pm 0.22$                               |
| $\Upsilon(2S) \rightarrow \chi_{b2}(1P)\gamma$ | 2.34  | $2.19 \pm 0.23$                                  | 1.84-1.93                        | $2.29 \pm 0.22$                               |
| $\chi_{b0}(2P) \rightarrow \Upsilon(2S)\gamma$ | 12.6  | $13.0 \pm 1.3$                                   | 10.6-11.4                        | -   |
| $\chi_{b1}(2P) \rightarrow \Upsilon(2S)\gamma$ | 17.1  | $16.3 \pm 1.7$                                   | 11.9-12.5                        | -   |
| $\chi_{b2}(2P) \rightarrow \Upsilon(2S)\gamma$ | 20.4  | $18.1 \pm 2.0$                                   | 12.9-13.1                        | -   |
| $\Upsilon(3S) \rightarrow \chi_{b0}(2P)\gamma$ | 1.44  | $1.05 \pm 0.14$                                  | 1.07-1.09                        | $1.20 \pm 0.16$                               |
| $\Upsilon(3S) \rightarrow \chi_{b1}(2P)\gamma$ | 2.38  | $2.05 \pm 0.24$                                  | 2.15-2.24                        | $2.56 \pm 0.34$                               |
| $\Upsilon(3S) \rightarrow \chi_{b2}(2P)\gamma$ | 2.53  | $2.35 \pm 0.25$                                  | 2.29-2.44                        | $2.66 \pm 0.41$                               |

Table : E1 decay rates for bottomonium. pNRQCD results compared to a potential model calculation (Grotch et al.) and and the current PDG values. LO denotes the result obtained without relativistic corrections, NLO indicates the result up to  $\mathcal{O}(v^2)$  neglecting color-octet effects in the weak-coupling regime and non-perturbative contributions to  $V_r^{(2)}$ . The error estimates give the generic size of one  $\mathcal{O}(v^2)$  correction as well as an estimate for the sum of all corrections at  $\mathcal{O}(v^3)$ . From Pietrulewicz.

| process                                    | $\Gamma_{\text{pNRQCD}}^{\text{LO}}/\text{keV}$ | $\Gamma_{\text{pNRQCD}}^{\text{NLO}}/\text{keV}$ | $\Gamma_{\text{mod}}/\text{keV}$ | $\Gamma_{\text{exp}}^{\text{PDG}}/\text{keV}$ |
|--|---|--|----------------------------------|---|
| $\chi_{c0}(1P) \rightarrow J/\psi\gamma$   | 199   | $158 \pm 60$                                     | 162-183                          | $122 \pm 11$                                  |
| $\chi_{c1}(1P) \rightarrow J/\psi\gamma$   | 421   | $302 \pm 126$                                    | 340-363                          | $296 \pm 22$                                  |
| $\chi_{c2}(1P) \rightarrow J/\psi\gamma$   | 568   | $415 \pm 170$                                    | 413-464                          | $386 \pm 27$                                  |
| $h_c(1P) \rightarrow \eta_c(1S)\gamma$     | 909   | $447 \pm 272$                                    | -                                | <600  |
| $\psi(2S) \rightarrow \chi_{c0}(1P)\gamma$ | 53.6  | $21.4 \pm 16.1$                                  | 26.0-40.3                        | $29.4 \pm 1.3$                                |
| $\psi(2S) \rightarrow \chi_{c1}(1P)\gamma$ | 45.2  | $30.7 \pm 13.6$                                  | 28.3-37.3                        | $28.0 \pm 1.5$                                |
| $\psi(2S) \rightarrow \chi_{c2}(1P)\gamma$ | 31.6  | $25.6 \pm 9.5$                                   | 17.5-22.7                        | $26.5 \pm 1.3$                                |
| $\eta_c(2S) \rightarrow h_c(1P)\gamma$     | 38.1  | $31.0 \pm 11.4$                                  | -                                | -   |

Table : E1 decay rates for charmonium. pNRQCD results at LO, NLO (including error estimate) compared to a potential model calculation (Grotch et al.) and the current PDG values. From Pietrulewicz.

## CONCLUSIONS

Precision is  $k_\gamma^3/m^2 \times \mathcal{O}(\alpha_s^2, v^2)$  for the allowed transitions. The convergence for the  $b\bar{b}$  ground state was quite good, and also quite reasonable for the  $c\bar{c}$  ground state and the  $b\bar{b}$  1P state.

$$\begin{aligned}\Gamma_{\Upsilon(1S) \rightarrow \eta_b(1S)\gamma} &= 15.18(51) \text{ eV}, \\ \Gamma_{J/\psi(1S) \rightarrow \eta_c(1S)\gamma} &= 2.12(40) \text{ keV}, \\ \Gamma_{h_b(1P) \rightarrow \chi_{b0}(1P)\gamma} &= 0.962(35) \text{ eV}, \\ \Gamma_{h_b(1P) \rightarrow \chi_{b1}(1P)\gamma} &= 8.99(55) \times 10^{-3} \text{ eV}, \\ \Gamma_{\chi_{b2}(1P) \rightarrow h_b(1P)\gamma} &= 0.118(6) \text{ eV}.\end{aligned}$$

Precision is  $k_\gamma^3/m^2 \times \mathcal{O}(v^4)$  for the forbidden transitions. Large logarithms associated with the heavy quark mass scale have also been resummed.

$$\Gamma_{\Upsilon(2S) \rightarrow \eta_b(1S)\gamma} = 6_{-06}^{+26} \text{ eV}.$$

This number is perfectly consistent with existing data.

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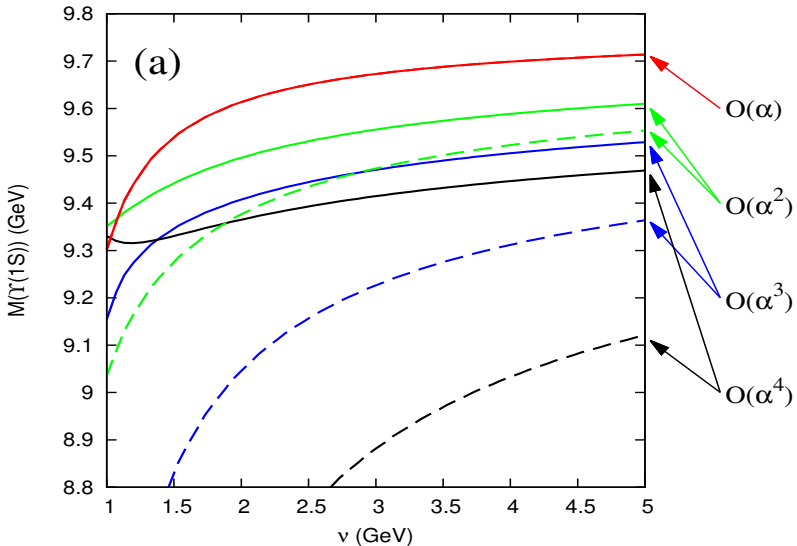


Figure :  $M_{10} = 2m_{b,RS'}(0.7 \text{ GeV}) + E_{10}$  using the static potential  $V_{RS'}^{(N)}$  at different orders in perturbation theory:  $N = 0, 1, 2, 3$ . Dashed lines with  $\nu_f = 0$ . Continuous lines with  $\nu_f = 0.7 \text{ GeV}$ . In both cases  $\nu_r = \infty \text{ GeV}$ .

$$V_{RS'}^{(N)}(r) = \begin{cases} (V^{(N)} + 2\delta m_{RS'}^{(N)})|_{\nu=\nu} \equiv \sum_{n=0}^N V_{RS',n} \alpha_s^{n+1}(\nu) & \text{if } r > \nu_r^{-1} \\ (V^{(N)} + 2\delta m_{RS'}^{(N)})|_{\nu=1/r} \equiv \sum_{n=0}^N V_{RS',n} \alpha_s^{n+1}(1/r) & \text{if } r < \nu_r^{-1}. \end{cases}$$

$$\delta m_X^{(N)}(\nu_f) = \nu_f \sum_{n=0}^N \delta m_X^{(n)}\left(\frac{\nu_f}{\nu}\right) \alpha_s^{n+1}(\nu)$$